

THE SEMICLASSICAL EINSTEIN-KLEIN-GORDON SYSTEM: ASYMPTOTIC ANALYSIS OF MINKOWSKI SPACETIME

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Einstein 1915: *Gravitational Interactions* $\longleftrightarrow$ *Lorentzian Manifolds*

$$\begin{cases} G_{ab} = \kappa T_{ab} \\ (\Box_g - m^2)\varphi = 0 \end{cases}$$

GEOMETRY $\iff$ Einstein tensor $G_{ab} := Ric_{ab} + g_{ab}(\Lambda - \frac{1}{2}R)$

MATTER $\iff$ stress-energy tensor T_{ab} of classical matter $\varphi : \mathcal{M} \rightarrow \mathbb{R}$

CONSTRAINTS: diffeomorphism invariance $\longleftrightarrow$ constructed bianchi identity

Baldeschi, Gelmini, Ruffini 1983: *Dark Matter* $\longleftrightarrow$ *classical Klein-Gordon field*

The Einstein-Klein-Gordon system is linearly stable:

Using a Taylor expansion $g_{ab} = \eta_{ab} + \epsilon h_{ab} + O(\epsilon^2)$ $\varphi = 0 + \epsilon \phi + O(\epsilon^2)$

$$\begin{cases} \Box h_{ab} = 0 & \text{gravitational wave equation} \\ (\delta h)_b := \eta^{ab} \partial_a h_{ab} = 0 & \text{harmonic gauge} \\ (\rho_\Sigma h)_{ab} = \gamma \in \ker K_\Sigma^\dagger := \ker(\rho_\Sigma \delta U) & \text{initial conditions} \end{cases}$$

$$\begin{cases} (\Box_\eta - m^2)\phi = 0 & \text{Klein-Gordon equation} \\ \rho_\Sigma \phi = f & \text{initial conditions} \end{cases}$$

WHAT HAPPEN FOR A QUANTIZED KLEIN-GORDON FIELD?

Møller 1962 & Rosenfeld 1963: *Classical Gravity* $\longleftrightarrow$ *Quantum Matter*

$$\begin{cases} G_{ab} = \kappa \langle :T_{ab}: \rangle_{\omega_2} \\ (\mathcal{P} \otimes 1) \omega_2 = (1 \otimes \mathcal{P}) \omega_2 = 0 \end{cases}$$

QUANTUM MATTER $\iff$ expectation value of renormalized quantum field $\langle :T_{ab}: \rangle_{\omega}$

CONSTRAINTS (for the system to be mathematically well-defined)

(1) $(\mathcal{M}, g_{ab})$ is *globally hyperbolic*, i.e. isometric to $(\mathbb{R} \times \Sigma, -\beta^2 dt^2 + \gamma_t)$

- (i) $\beta \in C^\infty(\mathcal{M}, (0, \infty))$
- (ii) γ_t is a Riemannian metric on Σ depending smoothly on $t \in \mathbb{R}$
- (iii) all sets $\{t_0\} \times \Sigma$ are Cauchy hypersurfaces in $\mathcal{M}$

(2) ω_2 is a bidistribution associated to an *Hadamard state*, i.e.

- (i) $\omega_2(f, f') - \omega_2(f', f) = i\mathcal{G}_{\mathcal{P}}(f, f')$ where $\ker \mathcal{P} \stackrel{\mathcal{G}_{\mathcal{P}}}{\simeq} C_c^\infty(\mathcal{M})$
- (ii) $|\mathcal{G}_{\mathcal{P}}(f, f')|^2 \leq 4\omega_2(f, f)\omega_2(f', f')$
- (iii) $\omega_2(f, f) > 0$ for all $f \notin \text{ran}(\mathcal{P})$
- (iv) $\text{WF}(\omega_2) = \{(x, k_x; y, -k_y) \in (T^*\mathcal{M} \times T^*\mathcal{M}) \setminus \{0\} \mid (x, k_x) \sim_{\parallel} (y, k_y), k_x \triangleright 0\}$

GOAL: STUDY (IN)STABILITY OF THE EINSTEIN-KLEIN-GORDON SYSTEM
LINEARIZED AROUND MINKOWSKI - POINCARÉ VACUUM

- (I) **The Quantization of a Klein-Gordon Field**
- (II) **The Formulation of the Cauchy Problem**
 - (i) the forcing problem for the linearized semiclassical system
 - (ii) decoupling the system via the Møller operator
 - (iii) tensor decomposition in the harmonic gauge
- (III) **Analysis of the Nonlocal Semiclassical System**
 - (i) the well-posedness of the forcing problem
 - (ii) asymptotic analysis of the solutions
- (IV) **Summary and Future Outlook**

Based on

“The Semiclassical Einstein-Klein-Gordon System” – arXiv:2604.01047

S. Galanda (York) P. Meda (Trento) N. Pinamonti (Genoa) G. Schmid (Regensburg)

ALGEBRAIC APPROACH TO QUANTUM FIELD THEORY

(1) *algebra of observables* $\mathcal{A}$ encoding causal structure, dynamics, and CCR

(2) *Hadamard state* $\omega: \mathcal{A} \rightarrow \mathbb{C}$ which is positive, normalized and **Hadamard**

$$(i) \quad \omega_2(f, f') - \omega_2(f', f) = i\mathcal{G}_{\mathcal{P}}(f, f') \quad \longleftrightarrow \quad \text{CCR}$$

$$(ii) \quad |\mathcal{G}_{\mathcal{P}}(f, f')|^2 \leq 4\omega_2(f, f)\omega_2(f', f') \quad \longleftrightarrow \quad \text{positivity}$$

$$(iii) \quad \omega_2(f, f) > 0 \quad \text{for all } f \notin \text{ran}(\mathcal{P})$$

$$(iv) \quad \text{WF}(\omega_2) = \{(x, k_x; y, -k_y) \in (T^*\mathcal{M} \times T^*\mathcal{M}) \setminus \{0\} \mid (x, k_x) \sim_{\parallel} (y, k_y), k_x \triangleright 0\}$$

Gérard and Wrochna 2012: Hadamard state $\iff b \in \Psi^1(\Sigma)$ such that

$$\mathcal{P} - (\partial_t + ib)(\partial_t - ib) = r_{-\infty}^{\pm} \in C_b^{-\infty}(\mathbb{R}, \Psi^{-\infty}(\Sigma_0))$$

For **ultrastatic spacetime**, i.e. $g_{ab} = -dt^2 + \gamma_{ij}$ and (Σ, γ_{ij}) is complete, being **Hadamard** is equivalent to solve the initial value problem for $\omega_2 \in D'(\mathcal{M} \times \mathcal{M})$

$$\begin{cases} (\mathcal{P} \otimes 1) \omega_2 = 0 \\ (1 \otimes \mathcal{P}) \omega_2 = 0 \\ \rho_{\Sigma} \omega_2 = \frac{1}{2} \begin{pmatrix} \sqrt{-\Delta_{\gamma_{ij}} + m^2}^{-1} & i\delta \\ -i\delta & \sqrt{-\Delta_{\gamma_{ij}} + m^2} \end{pmatrix} \end{cases}$$

Given $\omega : \mathcal{A} \rightarrow \mathbb{C} \xRightarrow{\text{GNS}} \pi : \mathcal{A} \rightarrow \mathcal{L}(\mathcal{H})$ operator-valued distributions but for $\phi : \mathcal{M} \rightarrow \mathbb{R}$

$$T_{ab} = \frac{1}{2} \nabla_a \nabla_b \phi^2 - \phi \nabla_a \nabla_b \phi - \frac{1}{2} g_{ab} \phi \square_g \phi - \frac{1}{2} g_{ab} \left(\frac{1}{2} \square_g \phi^2 + m^2 \phi^2 \right) + \xi (G_{ab} - \nabla_a \nabla_b + g_{ab} \square_g) \phi^2$$

involves pointwise products of field distributions $\implies$ general ill-defined

Brunetti, Fredenhagen and Kohler 1996: *Hadamard parametrix* $\mathcal{H} \in D'(\mathcal{M} \times \mathcal{M})$

$\omega_2 - \mathcal{H} \in C^\infty(\mathcal{M} \times \mathcal{M}) \implies$ well-defined restriction to $\text{diag}(\mathcal{M} \times \mathcal{M})$

Hollands and Wald 2001: *axiomatic regularisation scheme* encoding

- locality • covariance • conservation • analytic scaling behaviour

$$\langle : T_{ab} : \rangle_\omega = \lim_{y \rightarrow x} \mathcal{D}_{ab}(\omega_2 - \mathcal{H}) + \frac{A}{3} g_{ab} + \alpha_1 m^4 g_{ab} + \alpha_2 m^2 G_{ab} + \alpha_3 J_{ab} + \alpha_4 I_{ab}$$

- $\mathcal{D}_{ab}$ is a bidifferential operator derived from the classical expression for T_{ab}
- A is a coefficient depending on ξ , m , $R_{acbd} R^{acbd}$, $R_{ab} R^{ab}$, R and $\square_g R$
- $\alpha_i \in \mathbb{R}$ are *renormalisation constants*
- I_{ab} and J_{ab} geometric tensors depending on R_{ab} , $\square_g R_{ab}$, R and $\square_g R$

The semiclassical EKG system coupled ω_2 and g_{ab} and require two constraints

- (I) ω_2 is a Hadamard state;
- (II) $(\mathcal{M}, g_{ab})$ is globally hyperbolic.

PROBLEM: initial data for $\omega_2 \Leftrightarrow$ knowing g_{ab} in a neighbourhood of Σ

OUR PROPOSAL: The Cauchy problem will be a **FORCING PROBLEM**

TOY MODEL: 1st-order SHS $\mathcal{S} := \partial_t - \partial_x : C^\infty(\mathbb{R}^2) \rightarrow C^\infty(\mathbb{R}^2)$

using $\chi(t) := \begin{cases} 1 & t \geq 0 \\ 0 & t \leq -\epsilon \end{cases} \in C^\infty(\mathbb{R})$ and $(\mathcal{S} u)\chi = \mathcal{S}(\chi u) - (\partial_t \chi)u = 0$

$$(IVP) \quad \begin{cases} \mathcal{S} u = 0 \\ u|_{\{t=0\}} = u_0 \end{cases} \quad \xLeftrightarrow{t \geq 0} \quad (FP) \quad \begin{cases} \mathcal{S} \tilde{v} = f \in C_{pc}^\infty(\mathcal{M}) \\ \tilde{v} \in C_{pc}^\infty(\mathcal{M}) \end{cases}$$

The solution of (IVP) and the one of (FP) agrees for $t \geq 0$

LINEARIZED SEMICLASSICAL EINSTEIN-KLEIN-GORDON SYSTEM

We use a Taylor expansion $g_{ab} = \eta_{ab} + \epsilon h_{ab} + O(\epsilon^2)$ $\omega_2 = \omega_0 + \epsilon \omega_1 + O(\epsilon^2)$

(0) (η_{ab}, ω_0) solves the semiclassical EKG system $\iff$ finding α_1 s.t.

$$\langle :T_{ab}[\eta_{ab}]: \rangle_{\omega_0} = \lim_{y \rightarrow x} \mathcal{D}_{ab}(\omega_0 - \mathcal{H}) + \alpha_1 m^4 \eta_{ab} = 0$$

(1) the Cauchy Problem for the linearized semiclassical EKG system is

$$\left\{ \begin{array}{l} G_{ab}^{(1)}[h_{ab}] = \kappa \frac{d}{d\lambda} \left(\langle :T_{ab}[\eta_{ab} + \lambda h_{ab}]: \rangle_{[\omega_0 + \lambda \omega_1]} \right) \Big|_{\lambda=0} + S_{ab} \in C_{pc}^\infty(\mathcal{M} \times \mathcal{M}) \\ \frac{d}{d\lambda} \left((\mathcal{P}_{[\eta_{ab} + \lambda h_{ab}]} \otimes 1)(\omega_0 + \lambda \omega_1) \right) \Big|_{\lambda=0} = \Xi \in \Gamma_{pc}(\otimes_s^2 T^*\mathcal{M}) \\ \frac{d}{d\lambda} \left((1 \otimes \mathcal{P}_{[\eta_{ab} + \lambda h_{ab}]})(\omega_0 + \lambda \omega_1) \right) \Big|_{\lambda=0} = \Xi \in \Gamma_{pc}(\otimes_s^2 T^*\mathcal{M}) \\ \omega \in D'_{pc}(\mathcal{M} \times \mathcal{M}) \\ h_{ab} \in \Gamma_{pc}(\otimes_s^2 T^*\mathcal{M}) \end{array} \right.$$

PROBLEM: The linearized semiclassical EKG system is not decoupled

Since $g_{ab} \simeq \eta_{ab} + \epsilon h_{ab}$, the algebras of observables are $*$ -isomorphic

$$\begin{array}{ccc} \mathcal{R} : \mathcal{A}_{g_{ab}} & \xrightarrow{\simeq} & \mathcal{A}_{\eta_{ab}} \\ & \searrow \omega_{g_{ab}} := \mathcal{R}^* \omega_{\eta_{ab}} & \downarrow \omega_{\eta_{ab}} = \omega_{\mathbf{0}} + \tilde{\omega} \\ & & \mathbb{C} \end{array}$$

and the dynamics are intertwined $\mathcal{P}_{[g_{ab}]} \omega_{g_{ab}} = \mathcal{P}_{[g_{ab}]} \mathcal{R}^* \omega_{\eta_{ab}} = \mathcal{P}_{[\tilde{\eta}_{ab}]} \omega_{\eta_{ab}}$

So we can decoupled linearized EKG equations

$$\begin{cases} G_{ab}^{(1)}[h_{ab}] = \kappa \langle (: \mathcal{R}(T_{ab}[g_{ab}]) :)_{[\omega_{\eta_{ab}}]}^{(1)} \rangle = S_{ab} \\ h_{ab} \in \Gamma_{pc}(\otimes_s^2 T^* \mathcal{M}) \end{cases} \quad \begin{matrix} \nwarrow \\ \swarrow \end{matrix}$$

$$\begin{cases} (\mathcal{P}_{[\eta_{ab}]} \otimes 1) \omega_{\eta_{ab}} = \Xi' & \text{nonlocal operator} \\ (1 \otimes \mathcal{P}_{[\eta_{ab}]}) \omega_{\eta_{ab}} = \Xi' & \leftarrow \text{Cauchy problem under control} \\ \omega \in D'_{pc}(\mathcal{M} \times \mathcal{M}) \end{cases}$$

DEFINITION: Cauchy problem for the *linearized semiclassical Einstein equations*

$$\begin{cases} G_{ab}^{(1)}[h_{ab}] + \kappa \langle N_{ab}[h_{ab}] \rangle_{\omega_{\mathbf{0}}} - \kappa \langle : T_{ab}[h_{ab}] : \rangle_{\omega_{\mathbf{0}}} = S'_{ab} := S_{ab} - \kappa \langle : T_{ab}[\eta_{ab}] : \rangle_{\tilde{\omega}} \\ h_{ab} \in \Gamma_{pc}(\otimes_s^2 T^* \mathcal{M}) \end{cases}$$

The semiclassical Einstein equations inherit a *linear gauge symmetry*

$$h_{ab} \mapsto \tilde{h}_{ab} := h_{ab} + \partial_a X_b + \partial_b X_a - \eta_{ab} \partial^c X_c$$

THEOREM: For any $h_{ab} \in \Gamma_{\text{pc}}(\otimes_s^2 T^* \mathbb{R}^n)$ exists unique $X \in C_{\text{pc}}^\infty(T^* \mathbb{R}^n)$

$$\partial^a \tilde{h}_{ab} = 0 \quad (\text{de Donder gauge}) \quad \tilde{h}_{ab} = \tilde{h}_{ab}^S + \tilde{h}_{ab}^{\text{TT}} \quad (\text{tensor decomposition})$$

where $\tilde{h}_{ab}^{\text{TT}} \in \Gamma_{\text{pc}}(\otimes_s^2 T^* \mathbb{R}^n)$ with $\partial^a \tilde{h}_{ab}^{\text{TT}} = 0$ and $\eta^{ab} \tilde{h}_{ab}^{\text{TT}} = 0$

$$\tilde{h}_{ab}^S = \frac{1}{n-1} (\eta_{ab} \tilde{h}_c^c - \partial_a \partial_b G_{\text{ret}}(h))$$

Notice that $\partial^a \tilde{h}_{ab} = 0 \iff \partial^a h_{ab} = \square X_b \implies$ complete gauge fixing!

Forcing Problem for the linearized semiclassical Einstein equations

$$\begin{cases} -\frac{1}{2\kappa} \square h_{ab}^S + \frac{1}{4} \mathcal{K}_0[\mathcal{S} h_{ab}^S] - \frac{m^4}{64\pi^2} h_{ab}^S + \frac{1}{2} \tilde{\alpha}_2^S m^2 \square h_{ab}^S - \tilde{\alpha}_3^S \square \square h_{ab}^S = S_{ab}^S \in \Gamma_c(\otimes_s^2 T^* \mathcal{M}) \\ \partial^a h_{ab}^S = 0 \end{cases}$$

$$\begin{cases} -\frac{1}{2\kappa} \square h_{ab}^{\text{TT}} - \frac{1}{2} \mathcal{K}_0[\mathcal{T} h_{ab}^{\text{TT}}] + \frac{1}{2} \tilde{\alpha}_2^{\text{TT}} m^2 \square h_{ab}^{\text{TT}} - \tilde{\alpha}_4^{\text{TT}} \square \square h_{ab}^{\text{TT}} = S_{ab}^{\text{TT}} \in \Gamma_c(\otimes_s^2 T^* \mathcal{M}) \\ \partial^a h_{ab}^{\text{TT}} = 0 \end{cases}$$

where $\mathcal{K}_0(-) = \frac{1}{16\pi^2} \square \int_{m^2}^\infty \rho(M) G_{\text{ret}}^M(-) dM \quad \mathcal{S}, \mathcal{T} \in \text{Diff}_{\text{hyp}}^2(\otimes_s^2 T^* \mathcal{M})$

TOY MODEL for the Semiclassical Einstein Equations with $\phi \in C_{\text{pc}}^\infty(\mathcal{M})$

$$\mathcal{K}_0[(\square - a_1)(\square - a_2)\phi] + \underbrace{\sum_{j=0}^2 b_j \square^j \phi}_{\text{order 4}} = S$$

PROPOSITION $\iota_{\mathcal{M}} : \mathcal{M} \hookrightarrow \mathcal{M}_5 := \mathcal{M} \times \mathbb{R}$ and we consider

$$\tilde{\mathcal{G}}_{\text{ret}}^{4m^2} : C_{\text{pc}}^\infty(\mathcal{M}) \otimes L^2(\mathbb{R}) \rightarrow C_{\text{pc}}^\infty(\mathcal{M}) \otimes C^0(\mathbb{R})$$

Then, for every $\phi \in C_{\text{pc}}^\infty(\mathcal{M})$, we have that

$$\mathcal{K}_0(\phi) := \frac{1}{16\pi^2} \square \int_{m^2}^\infty \rho(M) \mathcal{G}_{\text{ret}}^M(\phi) dM = \iota_{\mathcal{M}}^* \tilde{\mathcal{G}}_{\text{ret}}^{4m^2}(\square\phi \otimes \check{\rho}) = \tilde{\mathcal{G}}_{\text{ret}}^{4m^2}(\square\phi \otimes \check{\rho})|_{z=0}$$

where $\check{\rho}(z) = \int_{\mathbb{R}} \rho(p_z^2 + 4m^2) |p_z| e^{ip_z z} dp_z$.

If ς is positive, square-integrable on $[4m^2, \infty)$ and Hölder continuous, then

$$C_{\text{pc}}^\infty(\mathcal{M}) \ni f \mapsto \mathcal{G}_\varsigma(f) := \iota_{\mathcal{M}}^* \tilde{\mathcal{G}}_{\text{ret}}^{4m^2}(f \otimes \varsigma) \quad \text{is invertible}$$

$$\text{TOY MODEL} \Leftrightarrow \underbrace{\iota_{\mathcal{M}}^* \tilde{\mathcal{G}}_{\text{ret}}^{4m^2}}_{\text{order -2}} \left(\underbrace{(\square(\square - a_1)(\square - a_2)\phi \otimes \check{\rho})}_{\text{order 6}} \right) + \underbrace{\sum_{j=0}^2 b_j \square^j \phi}_{\text{order 4}} = S$$

NOTICE: inverting $\iota_{\mathcal{M}}^* \tilde{\mathcal{G}}_{\text{ret}}^{4m^2}$ does not help at this stage

We start with rewriting the TOY MODEL in the following form

$$\iota_{\mathcal{M}}^* \tilde{\mathbf{G}}_{\text{ret}}^{4m^2} \left((\square(\square - a_1)(\square - a_2)\phi \otimes \check{\rho} + \sum_{j=0}^2 (\tilde{\square} - 4m^2)(b_j \square^j \phi \otimes \mathfrak{h}_j) \right) = S$$

where $\tilde{\square}$ is the d'Alembert operator on $\mathcal{M}_5$, and with a direct computation

$$\iota_{\mathcal{M}}^* \tilde{\mathbf{G}}_{\text{ret}}^{4m^2} \left((\square(\square - a_1)(\square - a_2)\phi \otimes \check{\rho} + \sum_{j=0}^2 ((\square - 4m^2)b_j \square^j \phi \otimes \mathfrak{h}_j) + \sum_{j=0}^2 b_j \square^j \phi \otimes \partial_z^2 \mathfrak{h}_j \right) = S$$

If $\mathfrak{h}_j$ satisfy the following ODEs for some constants β_i and with $\mathfrak{h}_j(0) = 1$

$$\begin{cases} b_2 \partial_z^2 \mathfrak{h}_2 - 4m^2 b_2 \mathfrak{h}_2 + b_1 \mathfrak{h}_1 - (a_1 + a_2) \check{\rho} = \beta_2 (\check{\rho} + b_2 \mathfrak{h}_2) \\ b_1 \partial_z^2 \mathfrak{h}_1 - 4m^2 b_1 \mathfrak{h}_1 + b_0 \mathfrak{h}_0 + (a_1 a_2) \check{\rho} = \beta_1 (\check{\rho} + b_2 \mathfrak{h}_2) \\ b_0 \partial_z^2 \mathfrak{h}_0 - 4m^2 b_0 \mathfrak{h}_0 = \beta_0 (\check{\rho} + b_2 \mathfrak{h}_2). \end{cases} \quad (1)$$

then the TOY MODEL simplifies as

$$\iota_{\mathcal{M}}^* \tilde{\mathbf{G}}_{\text{ret}} ((\square^3 \phi + \beta_2 \square^2 \phi + \beta_1 \square \phi + \beta_0 \phi) \otimes (\check{\rho} + b_2 \mathfrak{h}_2)) = S$$

if $\check{\rho} + b_2 \mathfrak{h}_2$ is positive and sufficiently regular, then the TOY MODEL reads

$$\square^3 \phi + \beta_2 \square^2 \phi + \beta_1 \square \phi + \beta_0 \phi = S_N$$

PROPOSITION: $\forall a_1, a_2 \leq 4m^2$ and $\forall \tilde{b}_0, \tilde{b}_1, \tilde{b}_2 \in \mathbb{R}$, there exist $b_0, b_1 \in \mathbb{R}$ s.t.

$$\begin{aligned}
 \iota_{\mathcal{M}}^* \tilde{G}_{\text{ret}}^{4m^2} (\square(\square - a_1)(\square - a_2)\phi \otimes \check{\rho}) + \sum_{j=0}^2 \tilde{b}_j \square^j \phi &= S \\
 \Updownarrow \\
 \underbrace{\square^3 \phi + \beta_2 \square^2 \phi + \beta_1 \square \phi + \beta_0 \phi}_{L\phi} + \mathcal{G}_{\zeta}^{-1}((\tilde{b}_1 - b_1)\square \phi + (\tilde{b}_0 - b_0)\phi) &= \mathcal{G}_{\zeta}^{-1}(S) \\
 \Updownarrow G_{L,\text{ret}} \\
 \phi + W_{\text{ret}}(\phi) &= S
 \end{aligned}$$

WE SOLVE PERTURBATIVELY

$$\phi_0 := S, \quad \phi_n := W_{\text{ret}}(\phi_{n-1}) \quad \text{for } n > 0.$$

THEOREM: The series defined by

$$\phi = \sum_{n \geq 0} (-1)^n \phi_n = \sum_{n \geq 1} (-1)^n W_{\text{ret}}^n(S)$$

converges absolutely in $C_{\text{pc}}^{\infty}(\mathcal{M})$

TOY MODEL $\mathcal{K}_0[(\square - a_1)(\square - a_2)\phi] + \sum_{j=0}^2 b_j \square^j \phi = S,$

$$\Updownarrow w^2 := s^2 + p^2$$

Fourier–Laplace $Q(w^2)(\mathcal{L}\mathcal{F}\phi(s, p)) = \mathcal{L}\mathcal{F}S(s, p)$

Hence, the solution of the **TOY MODEL** can be written as

$$\mathcal{L}\mathcal{F}(\phi)(s, p) = \frac{1}{Q(w^2)} \hat{S}(s, p)$$

Using Stieltjes-Perron inversion formula, we get

$$\phi = \sum_{\gamma \in \mathcal{Z}} \frac{1}{Q'(-\gamma)} G_{\text{ret}}^{\gamma}(S) + \int_{4m^2}^{\infty} \vartheta(M) G_{\text{ret}}^M(S) dM$$

where $\mathcal{Z}$ is the set of zeros of Q in $\mathbb{C} \setminus (-\infty, -4m^2)$.

THEOREM: (1) If $\mathcal{Z} \subset (-4m^2, 0) \iff \phi$ decays at least as $t^{-3/2}$ for large time

(2) If $\mathcal{Z} \subset (0, \infty) \iff \phi$ is exponentially growing

(S) b_2 is a free parameter which depends $\tilde{\alpha}_3^S$

$$a_1 = a_2 = \frac{2m^2}{6\xi - 1}, \quad b_0 = -\alpha_1^S \frac{4m^4}{6\left(\frac{1}{6} - \xi\right)^2}, \quad b_1 = -\frac{2}{\kappa} \frac{1}{6\left(\frac{1}{6} - \xi\right)^2}$$

(TT) b_2 is a free parameter proportional to $\tilde{\alpha}_4^{TT}$

$$a_1 = a_2 = 4m^2, \quad b_0 = 0, \quad b_1 = \frac{60}{\kappa}$$

MAIN RESULT: For any compactly supported Cauchy data, the solution of the semiclassical Einstein equations grows exponentially in time but bounded by

$$\tilde{h}_{ab} \leq e^{Ht} \eta_{ab}$$

with H an effective Hubble constant depending on m of the Klein-Gordon field



H match experiment if $m \simeq 7.8 \times 10^{-3}$ eV which is compatible with axions-particles

CONJECTURE: The de Sitter universe $(\mathbb{R}^4, g_{ab} := e^{Ht} \eta_{ab})$ is linearly stable

THANKS for your attention!